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Periodic orbits for an infinite family of classical superintegrable systems

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Abstract

We show that all bounded trajectories in the two-dimensional classical system with the potential $V(r, \varphi) = \omega^2 r^2 + \frac{\alpha k^2}{r^2 \cos^2 k \varphi} + \frac{\beta k^2}{r^2 \sin^2 k \varphi}$ are closed for all integer and rational values of k. The period is $T = \frac{\pi}{2\omega}$ and does not depend on k. This agrees with our earlier conjecture suggesting that the quantum version of this system is superintegrable.

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(Some figures in this article are in colour only in the electronic version)

1. Introduction

A recent article was devoted to an infinite family of quantum systems on a plane described by the Hamiltonian

$$H_k = -\partial_r^2 - \frac{1}{r}\partial_r - \frac{1}{r^2}\partial_{\varphi}^2 + \omega^2 r^2 + \frac{\alpha k^2}{r^2 \cos^2 k\varphi} + \frac{\beta k^2}{r^2 \sin^2 k\varphi},\tag{1}$$

where k is an arbitrary real number and (r, φ) are polar coordinates [1]. These systems were shown to be exactly solvable and integrable for all values of k. Integrability is assured by the existence of a second-order integral of motion

$$X_k = L_3^2 + \frac{\alpha k^2}{\cos^2 k\varphi} + \frac{\beta k^2}{\sin^2 k\varphi}, \qquad L_3 = -i\partial_{\varphi}, \tag{2}$$

where L_3 is the two-dimensional angular momentum.

The existence of the integral X_k implies the separation of variables in polar coordinates in the Schrodinger equation in quantum mechanics and also in the Hamilton–Jacobi equation in classical mechanics.

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Superintegrability in a two-dimensional system means that one further integral of motion exists. Let us call it Y_k .

It was conjectured in our previous article [1] that the system (1) is superintegrable for all integer values of k and that the additional integral of motion is of order 2k. This conjecture was proven for k = 1, 2, 3 and 4. If $\alpha = 0$ or $\beta = 0$, it was proven for k = 1 to 6 and 8. For k = 1, 2 and 3, the system (1) reduces to known superintegrable systems. For k = 1, this is one of the four systems found in [2, 3], for k = 2 it was found in [4] and for k = 3 in [5]. The three-body rational Calogero model [6] is a special case of (1) for k = 3.

In all cases (k = 1, 2, 3 and 4) the quantum Hamiltonian and the other quantum integrals of motion lie in the enveloping algebra of an underlying hidden Lie algebra [7] and this is presumably true for all integer values of k [1].

The purpose of this paper is to analyze the corresponding classical system with Hamiltonian

$$H_{k} = p_{r}^{2} + \frac{p_{\varphi}^{2}}{r^{2}} + \omega^{2}r^{2} + \frac{\alpha k^{2}}{r^{2}\cos^{2}k\varphi} + \frac{\beta k^{2}}{r^{2}\sin^{2}k\varphi}.$$
(3)

We show that all bounded trajectories for this system are closed and that the motion is periodic for all integer and rational values of k.

We recall that in general, in an *n*-dimensional space, maximal superintegrability means that the classical Hamiltonian allows 2n - 1 functionally independent integrals of motion (including the Hamiltonian) that are well-defined functions on the phase space. In classical mechanics maximal superintegrability implies that all bounded trajectories are closed and the motion is periodic [8]. The corresponding statement in quantum mechanics is that the existence of a non-Abelian algebra of integrals of motion implies that the energy levels of the system are degenerate. In the two-dimensional case it means that the energy depends on one quantum number and the wavefunction on two. This is indeed the case for Hamiltonian (1) for all rational values of k and is part of the justification of our conjecture [1].

In section 2 we integrate the Hamilton–Jacobi equation for the arbitrary values of k in polar coordinates, restricting to the case of bounded trajectories. In section 3 we obtain an explicit expression for the bounded trajectories in terms of Chebyshev polynomials and show that these trajectories are all closed for integer and rational values of k. The period in time is $T = \frac{\pi}{2\omega}$. For integer k with $k \ge 2$, up to k - 1 singular points can occur on the trajectories in which the tangent vector is not defined. The special case k = 1 is discussed in detail in section 4. This is the only case in which the Hamilton–Jacobi equation separates in a second coordinate system, namely in Cartesian coordinates. Section 5 is devoted to examples of trajectories illustrating all features presented in the earlier sections.

2. Integration of the Hamilton–Jacobi equation

2.1. The basic equations

The Hamilton-Jacobi equation for the considered system in polar coordinates has the form

$$H_k = \left(\frac{\partial S}{\partial r}\right)^2 + \frac{1}{r^2} \left(\frac{\partial S}{\partial \varphi}\right)^2 + \omega^2 r^2 + \frac{1}{r^2} \left(\frac{\alpha k^2}{\cos^2 k\varphi} + \frac{\beta k^2}{\sin^2 k\varphi}\right) = E.$$
 (4)

We set the integral of motion (2) equal to a constant A. Since coordinates in (4) separate we look for a solution in which the classical action has the form

$$S(r, \varphi, t) = S_1(r) + S_2(\varphi) - Et$$
 (5)

and obtain

$$r^2 \left(\frac{\partial S_1}{\partial r}\right)^2 + \omega^2 r^4 - Er^2 = -A \tag{6}$$

$$X_k = \left(\frac{\partial S_2}{\partial \varphi}\right)^2 + \frac{\alpha k^2}{\cos^2 k\varphi} + \frac{\beta k^2}{\sin^2 k\varphi} = A.$$
 (7)

Since we are interested in the trajectories, we do not need to solve for S_1 and S_2 , though that is obviously possible. Instead we follow the usual procedure [9] and calculate the derivatives of *S* with respect to *E* and *A*. We have

$$S_1 = \int \frac{\sqrt{-A + Er^2 - \omega^2 r^4}}{r} \,\mathrm{d}r\tag{8}$$

$$S_2 = \int \frac{\sqrt{A\cos^2 k\varphi \sin^2 k\varphi - \alpha k^2 \sin^2 k\varphi - \beta k^2 \cos^2 k\varphi}}{\cos k\varphi \sin k\varphi} \,\mathrm{d}\varphi,\tag{9}$$

and hence

$$\frac{\partial S}{\partial E} = \frac{\partial S_1}{\partial E} - t = \delta_1 \tag{10}$$

$$\frac{\partial S}{\partial A} = \frac{\partial S_1}{\partial A} + \frac{\partial S_2}{\partial A} = \delta_2,\tag{11}$$

where δ_1 and δ_2 are constants.

2.2. Integration of the radial part

From (8) and (10) we have

$$t + \delta_1 = \frac{1}{2} \int \frac{r \, \mathrm{d}r}{\sqrt{-A + Er^2 - \omega^2 r^4}}.$$
 (12)

The roots of the denominator are

$$(r_{1,2})^2 = \frac{1}{2\omega^2} (E \pm \sqrt{E^2 - 4\omega^2 A}).$$
 (13)

Bounded motion $r_1 \leqslant r \leqslant r_2$ is obtained for

$$E^2 - 4\omega^2 A \ge 0$$
 and $A \ge 0$. (14)

Performing the integration in (12) we obtain

$$r^{2} = \frac{1}{2\omega^{2}} (E + \sqrt{E^{2} - 4\omega^{2}A} \sin[4\omega(t + \delta_{1})]).$$
(15)

Thus, for $\omega > 0$, r^2 is indeed bounded and periodic with period $T = \frac{\pi}{2\omega}$ for all values of α and β as long as (14) is satisfied.

For $\omega = 0$, the integral (12) must be evaluated differently and we obtain

$$r^{2} = \frac{A + 4E^{2}(t + \delta_{1})^{2}}{E}.$$
(16)

We see that the trajectories for $\omega = 0$ are not bounded, independently of the values of alpha and beta.

The limiting case A = 0 corresponds to the harmonic oscillator for which all trajectories are bounded and periodic. In the other limiting case $E^2 - 4\omega^2 A = 0$, the integral (12) is not defined. Directly from (8) we obtain $\frac{\partial S_1}{\partial r} = 0$, $r = \frac{1}{\omega}\sqrt{\frac{E}{2}}$, i.e. the trajectories are circles.

2.3. Integration of the angular part

From (8) and (9) we obtain

$$\frac{\partial S_1}{\partial A} = -\frac{1}{2} \int \frac{\mathrm{d}r}{r\sqrt{-A + Er^2 - \omega^2 r^4}} \tag{17}$$

$$\frac{\partial S_2}{\partial A} = \frac{1}{2} \int \frac{\sin k\varphi \cos k\varphi \, d\varphi}{\sqrt{A \cos^2 k\varphi \sin^2 k\varphi - \alpha k^2 \sin^2 k\varphi - \beta k^2 \cos^2 k\varphi}}.$$
(18)

The integrals are readily evaluated to give

$$\frac{\partial S_1}{\partial A} = -\frac{1}{4\sqrt{A}} \arcsin\left[\frac{-2A + Er^2}{r^2\sqrt{E^2 - 4\omega^2 A}}\right]$$
(19)

$$\frac{\partial S_2}{\partial A} = -\frac{1}{4k\sqrt{A}} \arcsin\left[\frac{-2A\sin^2 k\varphi + A - (\alpha - \beta)k^2}{\sqrt{(A - \alpha k^2 + \beta k^2)^2 - 4k^2\beta A}}\right].$$
(20)

In addition to (14) we require

$$(A - \alpha k^2 + \beta k^2)^2 - 4k^2 \beta A > 0$$
 and $\beta > 0.$ (21)

A more symmetric way of writing this condition is

$$[A - (\alpha + \beta)k^2]^2 - 4\alpha\beta k^2 \ge 0.$$
⁽²²⁾

From the form of Hamiltonian (3) we see that the trajectories are restricted to the sectors

$$\frac{n\pi}{2k} < \varphi < \frac{(n+1)\pi}{2k}, \qquad n \in \mathbb{R}.$$
(23)

The integral (18) was evaluated using the substitution $z = \sin^2 k\varphi$. The denominator in the integral vanishes for the values

$$z_{1,2} = \frac{(A - \alpha k^2 + \beta k^2) \mp \sqrt{(A - \alpha k^2 + \beta k^2)^2 - 4A\beta k^2}}{2A}.$$
 (24)

Condition (21) is necessary to assure $z_1 < z < z_2$ which is equivalent to (22). A final condition is $0 \le z_{1,2} \le 1$ which implies

$$A + k^2(\beta - \alpha) \ge 0$$
 and $\alpha > 0.$ (25)

3. The bounded trajectories

3.1. Equation for trajectories for integer values of k

So far we have established that bounded trajectories for Hamiltonian (3) are obtained if the constants of the system satisfy

$$A > 0, \quad \alpha > 0, \quad \beta > 0, \quad \omega > 0, \quad A > k^{2}|\beta - \alpha|, [A - (\alpha + \beta)k^{2}]^{2} - 4\alpha\beta k^{2} > 0, \qquad E^{2} - 4\omega^{2}A > 0.$$
(26)

Now let us obtain an equation for the trajectories. From (11), (19) and (20) we have

$$k \arcsin R + \arcsin U_k = -4k\sqrt{A\delta_2},\tag{27}$$

where we have introduced

$$R = \frac{-2A + Er^2}{r^2 \sqrt{E^2 - 4\omega^2 A}} \quad \text{and} \quad U_k = \frac{-2A \sin^2 k\varphi + A - (\alpha - \beta)k^2}{\sqrt{[A - (\alpha + \beta)k^2]^2 - 4\alpha\beta k^2}},$$
 (28)

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for r^2 and $z = \sin^2 k\varphi$ in the intervals

$$\frac{1}{2\omega^2}(E - \sqrt{E^2 - 4\omega^2 A}) \leqslant r^2 \leqslant \frac{1}{2\omega^2}(E + \sqrt{E^2 - 4\omega^2 A})$$
(29)

$$z_1 \leqslant z \leqslant z_2, \tag{30}$$

where $z_{1,2}$ are as in (24).

We have

$$-1 \leqslant R \leqslant 1, \qquad -1 \leqslant U_k \leqslant 1. \tag{31}$$

Once *R* is known as a function of φ we obtain $r(\varphi)$ as

$$r(\varphi) = \sqrt{\frac{2A}{E - R\sqrt{E^2 - 4\omega^2 A}}}.$$

We rewrite (27) as

$$k \arccos R = -\arccos U_k + C, \tag{32}$$

with

$$C = \frac{(k+1)\pi}{2} + 4k\sqrt{A}\delta_2, \qquad 0 \leqslant C \leqslant 2\pi.$$
(33)

Equation (32) can be converted into an algebraic equation for *R* in term of $\sin^2 k\varphi$. We use the well-known formula [10] for the Chebyshev polynomials

$$T_k(x) = \cos(k \arccos x) = \frac{(x + i\sqrt{1 - x^2})^k + (x - i\sqrt{1 - x^2})^k}{2}$$
(34)

for $k \ge 1$ and obtain

$$T_k(R) = \frac{(R + i\sqrt{1 - R^2})^k + (R - i\sqrt{1 - R^2})^k}{2} = U_k \cos C \pm \sqrt{1 - U_k^2} \sin C.$$
 (35)

We see that we can restrict the values of *C* to $0 \le C \le \pi$ since the interval $\pi \le C \le 2\pi$ will give the same trajectories.

Let us write (35) as a polynomial equation with real coefficients for *R* (and thus for r^2). We do this separately for even and odd values of *k* and obtain

$$k = 2m:$$

$$T_{2m} = \sum_{l=0}^{m} \binom{2m}{2l} \sum_{j=0}^{l} (-1)^{j} \binom{l}{j} R^{2(m-j)} = U_{2m} \cos C \pm \sqrt{1 - U_{2m}^{2}} \sin C,$$

$$k = 2m \pm 1;$$
(36a)

$$k = 2m + 1;$$

$$T_{2m+1} = \sum_{l=0}^{m} {\binom{2m+1}{2l}} \sum_{j=0}^{l} (-1)^{j} {\binom{l}{j}} R^{2(m-j)+1} = U_{2m+1} \cos C \pm \sqrt{1 - U_{2m+1}^{2}} \sin C.$$
 (36b)

The crucial point is that (35) does not contain any multivalued functions. It follows that $R(\varphi)$ and ultimately $r(\varphi)$ is a periodic function with the same period τ as $\sin^2 k\varphi$ namely $\tau = \frac{\pi}{k}$ and

$$r(\varphi) = r\left(\varphi + \frac{\pi}{k}\right). \tag{37}$$

Hence, all bounded trajectories are closed and the motion is periodic.

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We stress that $\tau = \frac{\pi}{k}$ is the period of *r* as a function of φ . Both *r* and φ are periodic in time *t* with the same period, namely $T = \frac{\pi}{2\omega}$ (see (15)), and this period does not depend on *k* and coincides with that of the harmonic oscillator.

Alternative (and equivalent) equations for the trajectories can be obtained. If we solve (35) for U_k we obtain

$$U_k(\varphi) = T_k(R) \cos C \pm \sqrt{1 - T_k(R)^2} \sin C.$$
 (38)

Since we have $|U_k| \leq 1$, we can define η_k by putting

$$\cos(\eta_k \pm C) = U_k \tag{39}$$

and rewrite (35) as

$$T_k(R) = \cos \eta_k. \tag{40}$$

In section 4 we shall plot the bounded trajectories for low values of k. The equations for the trajectories then simplify and we have

$$k = 1$$
: $R = U_1 \cos C \pm \sqrt{1 - U_1^2} \sin C$ (41)

$$k = 2: \quad 2R^2 - 1 = U_2 \cos C \pm \sqrt{1 - U_2^2 \sin C}$$
(42)

$$k = 3: \quad 4R^3 - 3R = U_3 \cos C \pm \sqrt{1 - U_3^2 \sin C}$$
(43)

$$k = 4: \quad 8R^4 - 8R^2 + 1 = U_4 \cos C \pm \sqrt{1 - U_4^2 \sin C}. \tag{44}$$

The trajectories are always closed curves within the sectors (23). They can never actually reach the boundaries of the sectors, unless we have $\alpha = 0$ or $\beta = 0$ (or both). The trajectories degenerate into line segments for $U_k^2 = 1$ (see (35)). The length of the segment is determined by the limits (29) and (30) which must be imposed.

3.2. Singular points on the trajectories for k integer

Let us consider a curve given in polar coordinates. Singular points on the curve are those where the derivative $\frac{dr}{d\omega}$ is not defined. They satisfy the overdetermined system of equations:

$$F(r,\varphi) = 0, \qquad \frac{\partial F}{\partial r} = 0, \qquad \frac{\partial F}{\partial \varphi} = 0.$$
 (45)

In our case the equation of the curve is

$$F_k(r,\varphi) = T_k(R) - U_k \cos C \mp \sqrt{1 - U_k^2} \sin C = 0,$$
(46)

and the derivatives satisfy

$$\frac{\partial F_{2m}}{\partial r} = \frac{\partial T_{2m}}{\partial R} \frac{\mathrm{d}R}{\mathrm{d}r}$$

$$= 2 \sum_{l=0}^{m} {\binom{2m}{2l}} \sum_{j=0}^{l} (-1)^{j} {\binom{l}{j}} (m-j) R^{2(m-j)-1} \frac{\mathrm{d}R}{\mathrm{d}r}$$

$$\partial F_{2m+1} = \frac{\partial T_{2m+1}}{\partial T_{2m+1}} \mathrm{d}R$$
(47)

$$\frac{2m+1}{\partial r} = \frac{2m+1}{\partial R} \frac{1}{dr}$$
$$= \sum_{l=0}^{m} {\binom{2m+1}{2l}} \sum_{j=0}^{l} (-1)^{j} {\binom{l}{j}} [2(m-j)+1] R^{2(m-j)} \frac{dR}{dr}$$
(48)

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$$\frac{\partial F_k}{\partial \varphi} = \frac{-\cos C \sqrt{1 - U_k^2 \pm U_k \sin C}}{\sqrt{1 - U_k^2}} \frac{\mathrm{d}U_k}{\mathrm{d}\varphi}.$$
(49)

From equation (28) we see that $\frac{dR}{dr}$ vanishes only for $r \to \infty$ and $\frac{dU_k}{d\varphi}$ only for $\varphi = \frac{n\pi}{2k}$, $n = 0, \pm 1, \pm 2, \dots$ These values correspond to points that do not lie on a trajectory.

The singular points satisfy $U_k = \pm \cos C$ that is

$$\sin^{2} k\varphi = \frac{A + k^{2}(\beta - \alpha) \pm \sqrt{(A + k^{2}(\beta - \alpha))^{2} - 4A\beta k^{2} \cos C}}{2A},$$
 (50)

and the one of equation (49) for k even or odd respectively.

From (49) we see that $\frac{\partial F_k}{\partial r}$ is a polynomial of order k-1 and will hence have k-1 zeros, all of them real. These zeros are actually singular points if the corresponding values of r and φ lie on the trajectories (see (46)). This in turn depends on the constants A, C and E, i.e. on the initial conditions (for ω , α and β given). The statement thus is: the number n_0 of singular points on the trajectories satisfies

$$0 \leqslant n_0 \leqslant k - 1. \tag{51}$$

For small values of *k* these formulas give: For $k = 1 : \frac{\partial F_1}{\partial r} = \frac{dR}{dr} \neq 0$ $r_1 \leq r \leq r_2$. There are no singular points.

For k = 2: $\frac{\partial F_2}{\partial r} = 4R \frac{dR}{dr} = 0$. The possible singular point corresponds to $r = \sqrt{\frac{2A}{E}}$. For k = 3: $\frac{\partial F_3}{\partial r} = 3(4R^2 - 1)\frac{dR}{dr} = 0$. The two possible singular points are $R = \pm \frac{1}{2}$ i.e. $r^2 = \frac{4A}{2E \pm \sqrt{E^2 - 4\omega^2 A}}$. For k = 4: $\frac{\partial F_3}{\partial r} = 16R(2R^2 - 1)\frac{dR}{dr} = 0$. The three possible singular points are $R_1 = 0$ and $R_{2,3} = \pm \frac{1}{\sqrt{2}}$, i.e. $r = \sqrt{\frac{2A}{E}}$, $r^2 = \frac{2\sqrt{2}A}{\sqrt{2E \pm \sqrt{E^2 - 4\omega^2 A}}}$.

3.3. Equation for trajectories for rational values of k = m/n

In the last subsection, we obtained the equation for the trajectories for integer values of k. We have established that those trajectories for Hamiltonian (3) are bounded if the constants of the system satisfy conditions (26). Now let us obtain the trajectories when k is rational.

Here we set k = m/n for m, n integers. The bounded trajectories in this case are obtained under the same conditions on the constants of the system as those specified in the preceding subsection (26).

Equation (32) becomes

$$m \arccos R = -n \arccos U_k + \hat{C}.$$
⁽⁵²⁾

Using identity (34) we obtain

$$T_m(R) = T_n(U_k) \cos \tilde{C} \pm \sqrt{1 - T_n(U_k)^2} \sin \tilde{C}.$$
(53)

Following the same arguments as in subsection 3.1, we see that $R(\varphi)$ and $r(\varphi)$ are periodic functions of the angle φ of period $\tau = \frac{\pi}{k} = \frac{n\pi}{m}$. Again all bounded trajectories are closed and the motion is periodic.

In section 5 we will present the trajectories for the rational values k = 1/2, 1/3, 3/2.

3.4. Comments on the range of the parameters

The trajectories given by (35) or (36b) (see also (41)–(44)) and (53) depend on three quantities E, A and C. E and A are the values of the energy H_k (3) and the integral of motion X_k (7), respectively. The third integral Y_k , responsible for superintegrability has so far not figured explicitly and is related to the constant *C*. The quantities *A* and *E* are restricted by relations (26) in order for the trajectories to be bounded. We allow *C* to take values $0 \le C \le \pi$. The trajectories could also be calculated differently, using all three integrals of motion and just one of the Newton equations of motion. This may impose further conditions on the parameters involved. We discuss this in detail for k = 1 in section 4.

4. The special case k = 1

The bounded trajectories for k = 1 are given by (41) with R(r) and $U_1(\varphi)$ defined in (28). The case k = 1 differs from that of all other values of k by the fact that the additional integral of motion Y_k for k = 1 is second order in the momenta and hence also leads to the separation of variables in the Hamilton–Jacobi equation. Explicitly we have [2, 3]

$$Y_1 = E_1 = P_1^2 + \omega^2 x^2 + \frac{\alpha}{x^2},$$
(54)

so Y_1 is the energy related to motion in the *x* direction (equally well we could have put $Y_1 = E_2 = E - E_1$ or $Y_1 = E_1 - E_2$ where E_2 is the energy related to motion in the *y* direction). The Hamilton–Jacobi equation in this case is multiseparable. Namely it separates in polar coordinates (for any *k*) and Cartesian ones (for k = 1). For k = 1, it also separates in elliptic coordinates with an arbitrary focal distance of $0 < f < \infty$.

The equations of motion were solved in Cartesian coordinates [2, 3] and the solutions corresponding to bounded trajectories in the notations of the present paper are

$$x^{2} = \frac{E_{1}}{2\omega^{2}} + \frac{1}{2\omega^{2}}\sqrt{E_{1}^{2} - 4\alpha\omega^{2}}\sin[4\omega(t+C_{1})]$$
(55)

$$y^{2} = \frac{E_{2}}{2\omega^{2}} + \frac{1}{2\omega^{2}}\sqrt{E_{2}^{2} - 4\beta\omega^{2}}\sin[4\omega(t+C_{2})].$$
(56)

Above we have $E_1 + E_2 = E$ and C_1 , C_2 are the integration constants. From (55) and (56) we see that the trajectories are periodic. They lie inside a rectangle

$$\frac{E_1 - \sqrt{E_1 - 4\alpha\omega^2}}{2\omega^2} \leqslant x^2 \leqslant \frac{E_1 + \sqrt{E_1 - 4\alpha\omega^2}}{2\omega^2}$$
(57)

$$\frac{E_2 - \sqrt{E_2 - 4\beta\omega^2}}{2\omega^2} \leqslant y^2 \leqslant \frac{E_2 + \sqrt{E_2 - 4\beta\omega^2}}{2\omega^2}.$$
(58)

We see that we must have

$$2\omega\sqrt{\alpha} \leqslant E_1 \leqslant E - 2\omega\sqrt{\beta}. \tag{59}$$

Eliminating t from (55) and (56), we obtain an equation for the trajectories, namely

$$y^{2} = \frac{E_{2}}{2\omega^{2}} + \frac{\sqrt{E_{2}^{2} - 4\beta\omega^{2}}}{2\omega^{2}\sqrt{E_{1}^{2} - 4\alpha\omega^{2}}} [\cos S(-E_{1} + 2\omega^{2}x^{2})$$
(60)

$$\pm 2\omega \sin S \sqrt{-\alpha + E_1 x^2 - \omega^2 x^4}$$
(61)

for $0 \leq S = C_1 - C_2 \leq \pi$.

This equation is equivalent to a fourth-order polynomial equation involving the coordinates x and y. Viewing x as a parameter for the curve, we note that it must satisfy (57). The curve

(60) has four branches, each lying entirely within one quadrant of the (x, y) plane. For $\alpha > 0$ and $\beta > 0$, it can never reach the coordinate axes (see (57) and (58)).

In the limiting cases when sin S = 0, the closed curve (60) collapses into a segment of a quadric. For S = 0, this is a segment of a hyperbola and for $S = \pi$, a segment of an ellipse (always inside the rectangle (57) and (58)).

The trajectory (41) in polar coordinates can be made more explicit, namely

$$r^{2} = \frac{2A}{E - \sqrt{E^{2} - 4\omega^{2}A}(U\cos C \pm \sqrt{1 - U^{2}}\sin C)}.$$
 (62)

Again this corresponds to a fourth-order polynomial curve (in *r* and sin φ) that degenerates into a segment of a quadric for sin C = 0. Considering $z = \sin^2 \varphi$ as a parameter of the curve we must restrict to the interval (30).

The trajectories given by (60) and (62) must coincide. This gives a relationship between the 'Cartesian constants' { E_1 , E, S} and the 'polar constants' {A, E, C}. We calculate r(t)and $\varphi(t)$ from (55) and (56) and compare with r(t) from (15) and $\varphi(t)$ from (41). For given total energy E (and α , β , ω), we obtain

$$A = \frac{E_1 E_2 + 2(\alpha + \beta)\omega^2 - \sqrt{(E_1^2 - 4\alpha\omega^2)(E_2^2 - 4\beta\omega^2)\cos S}}{2\omega^2},$$
 (63)

and

$$\cos C = \frac{A(2E_1 - E) - E(\alpha - \beta)}{\sqrt{(A^2 + (\alpha - \beta)^2 - 2A(\alpha + \beta))(E^2 - 4\omega^2 A)}}.$$
(64)

The condition $|\cos C| \le 1$ is not necessarily automatically satisfied by the right-hand side of (64) and imposes further conditions. First of all, the denominator in (64) must be the square root of a positive number which implies

$$A \ge (\sqrt{\alpha} + \sqrt{\beta})^2. \tag{65}$$

Secondly, the equation $\cos^2 C = 1$ is a quadratic polynomial in A and E_1 :

$$\omega^2 A^2 + A(E_1^2 - E_1 E - 2(\alpha + \beta)\omega^2) + E^2 \alpha + E_1 E(-\alpha + \beta) + (\alpha - \beta)^2 \omega^2 = 0.$$
 (66)

An example of the corresponding curve in the (E_1, A) plane is given in figure 1. We have chosen E = 20, $\omega = 2$, $\alpha = 3$ and $\beta = 2$. The curve $\cos^2 C = 1$ lies inside the rectangle

$$2\omega\sqrt{\alpha} \leqslant E_1 \leqslant E - 2\alpha\sqrt{\omega}, \qquad (\sqrt{\alpha} + \sqrt{\beta})^2 \leqslant A \leqslant \frac{E}{4\omega^2}$$

We have $0 < \cos C < 1$ in region I, $-1 < \cos C < 0$ in II, $\cos C < -1$ in III and IV, and $\cos C > 1$ in V and VI.

The dashed curve represents $\cos C = 0$. At the intersection points of this curve and (66), we have $\cos C = 0/0$, i.e. $\cos C$ is undefined.

5. Examples of trajectories

In this section we present curves showing trajectories for specific choices of parameters ω , α and β and the integrals of motion *E*, *A* and *C*. We choose k = 1, 2, 3, 4, 1/2, 1/3, 3/2.

From all of the figures we see that the trajectories do indeed lie within the sectors (23) and never reach the sector boundaries. From formulas in sections 2 and 3 and from the figures we see that we can actually restrict to one sector only, e.g. the one with n = 0 in (23). The trajectories in all other sectors are obtained by reflections in the sector boundaries.



Figure 1. Regions of bounded trajectories in the (E_1, A) plane (I and II).



Figure 2. Examples of trajectories for k = 1.

The figures illustrate all features discussed in the previous sections. For k = 1, the trajectories have no singular points. The dependence on the constant *C* is quite spectacular. For all values of *C* the trajectories are fourth order curves. For $C = \frac{\pi}{2}$, they resemble ellipses. For C = 0 or $C = \pi$, the trajectories flatten out and finally degenerate into line segments (see figure 2). This happens for all values of *k*. For k = 2, (42) has two real roots. For each of



Figure 3. Examples of trajectories for k = 2.



Figure 4. Trajectories for the real root for k = 3.

them we present the trajectories in all eight sectors for just one value of *C* (see figure 3). We see that the chosen trajectories have one singular point. For k = 3, (43) has three roots but only one is real. A typical set of trajectories is on figure 4. For k = 4, (44) has four real roots. We choose just one of them and present the trajectories in the first quadrant of the (x, y) plane (see figure 5).

Trajectories for rational values of k, namely k = 1/2, 1/3, 3/2 are given in figures 6–8, respectively. They somewhat resemble Lissajous curves for an anisotropic oscillator. The symmetry with respect to reflection in the boundaries is obvious in figures 2(f) to 5.



Figure 5. Trajectories for the first real root in the first quadrant for k = 4.



Figure 6. Trajectories for k = 1/2.

For k = 1, we set E = 20, A = 12, $\omega = 2$, $\alpha = 3$ and $\beta = 2$. The trajectories are given on figure 2.

For k = 2, we set E = 50, A = 60, $\omega = 2$, $\alpha = 3$ and $\beta = 2$. The trajectories in the eight sectors for $C = \frac{\pi}{2}$ are shown in figure 3.

For k = 3, we set E = 50, A = 100, $\omega = 2$, $\alpha = 3$ and $\beta = 2$, and for k = 4, E = 16, $A = 15, \omega = 2, \alpha = \frac{1}{4}$ and $\beta = \frac{1}{8}$. The trajectories are shown in figures 4 and 5. For k = 1/2 and 1/3, we set $E = 20, A = 24, \omega = 2, \alpha = 3$ and $\beta = 2$ in (53). For

 $C = \frac{\pi}{2}$, the trajectories are given in figures 6 and 7. There are no singular points.

For k = 3/2, we set E = 15, A = 10, $\omega = 2$, $\alpha = 1/4$ and $\beta = 1/8$ in (53). We obtain three roots for r^2 . Two of them are complex conjugate to each other. The trajectories corresponding to the real root for $C = \frac{\pi}{2}$ are shown in figure 8. We observe two singular points.



Figure 7. Trajectories for k = 1/3.



Figure 8. Trajectories for k = 3/2.

6. Conclusion

The present paper, together with our previous one [1], provides a very strong indication that Hamiltonian (3) is superintegrable for all integer values of k in classical and in quantum mechanics. In classical mechanics, the results of section 3.3 show that this is also true for rational values of k. Indeed, we have shown in [1] that the system (1) is exactly solvable for

all values of k thus supporting a previous conjecture [11] that all superintegrable systems are exactly solvable. A spectacular result is that the period $T = \frac{\pi}{2\omega}$ is not only independent of k but is the same as for the harmonic oscillator.

We mention that Hamiltonian (3) can be interpreted as representing a three-body system on a line, once the motion of the center of mass is factored out. The same classical system was studied in [12] for $\omega = 0$ and $\alpha = 0$. In this case the trajectories are never bounded (see (26)) and the motion cannot be periodic. A third integral of motion can still exist and the authors suggest a possible form of the additional integral for *k* odd [12].

A rigorous proof of superintegrability requires the explicit construction of the third integral of motion for all integer and rational values of k and possibly even for any $k > 0, k \in \mathbb{R}$. This is left for a future study.

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